

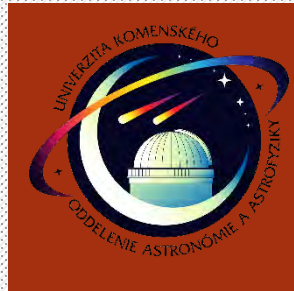


Comenius University

RNDr. Roman Nagy, PhD.

Leonard Kornoš

FMFI UK



AGO Modra



Students and colleagues



Students and colleagues



V Akademickom roku 2021/2022

- 13 magisterských študentov
- 9 doktorandov

Astronomy at Comenius University



- fields of research:
 - interplanetary matter
 - meteor astronomy
 - galactic astrophysics
 - space debris

Interplanetary matter



- astrometry and photometry of asteroids and comets
 - improvement of orbits
- dynamics and physical properties of asteroids (rotation rates, Yarkovski a YORP effects,...)
- theoretical research on gravitational motion and non-gravitational effects

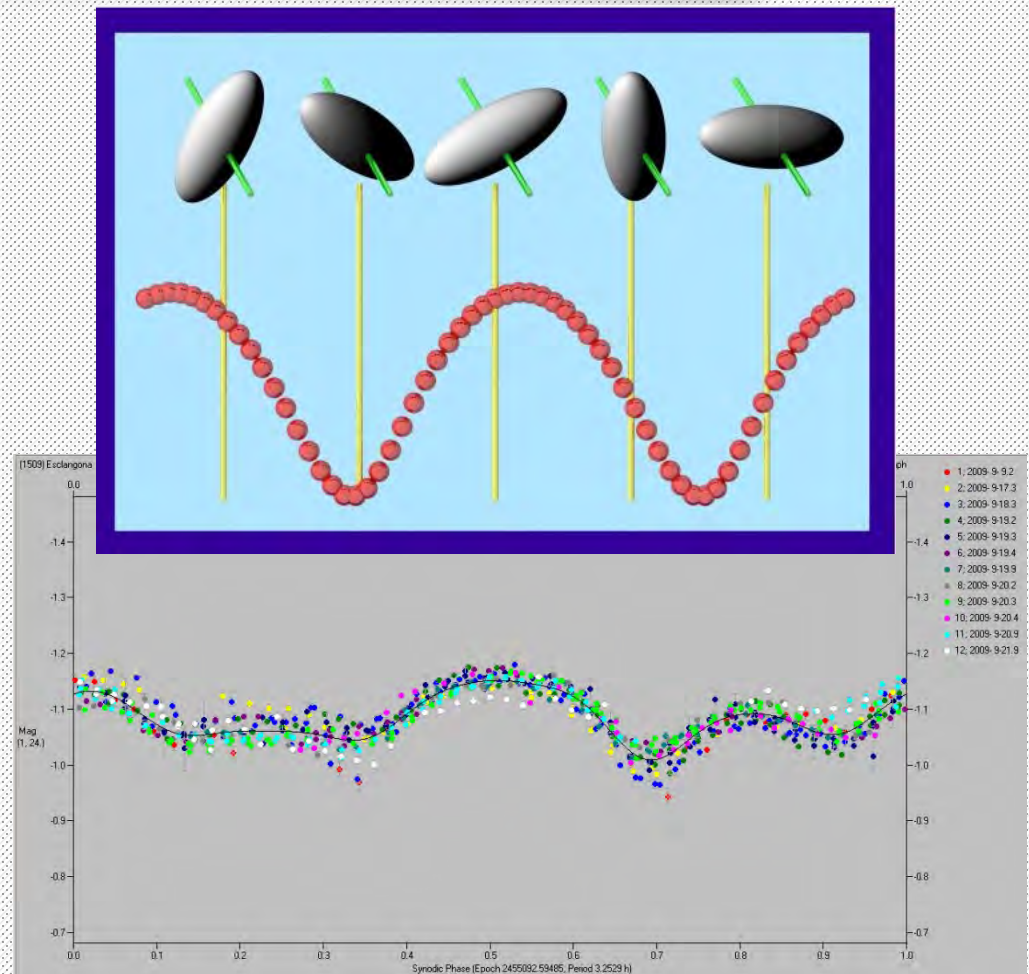
60 cm telescope



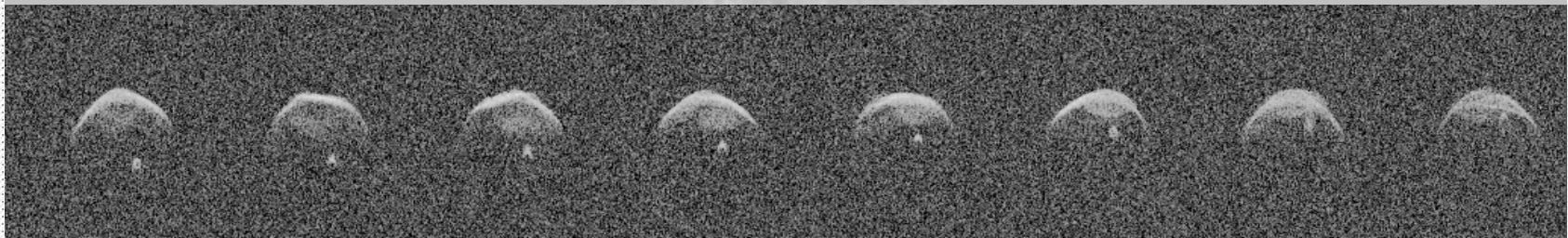
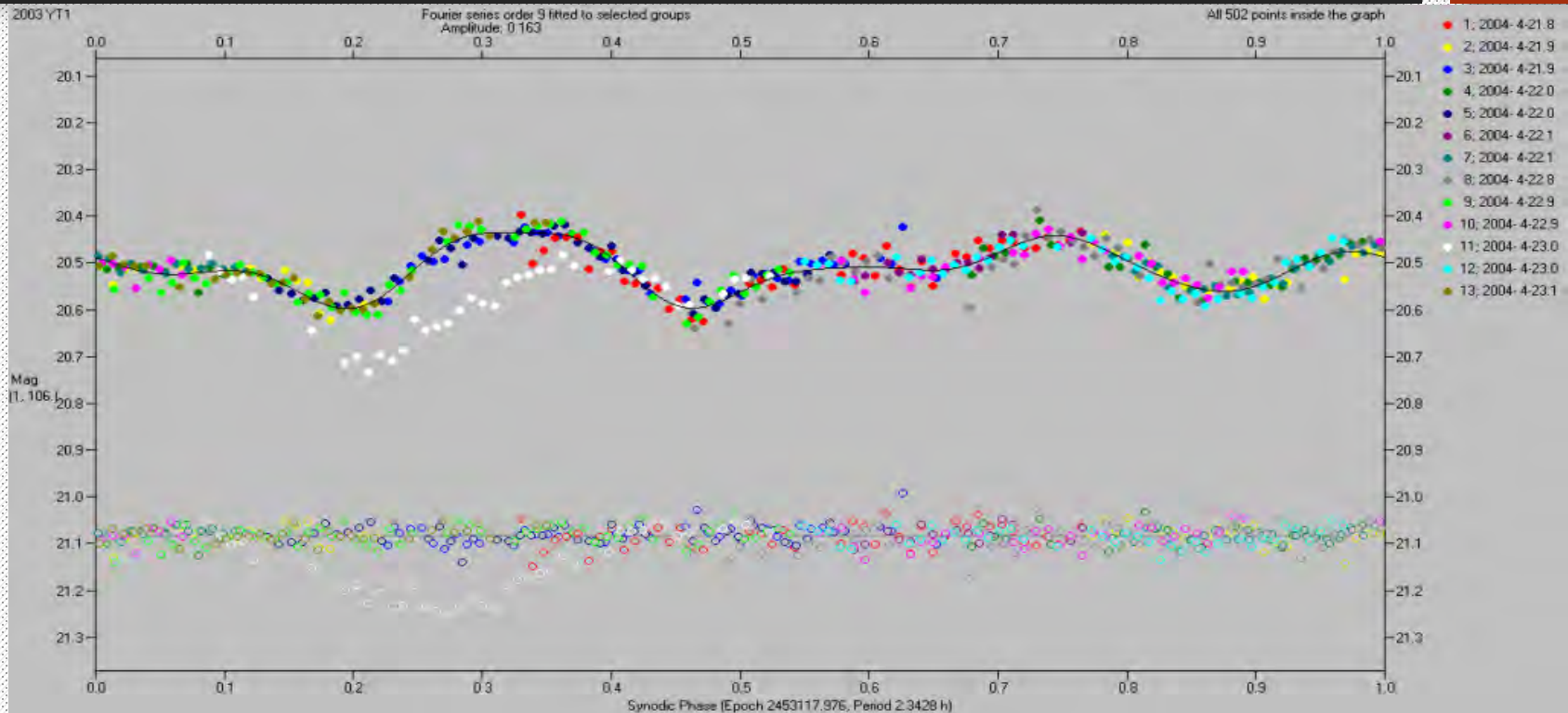
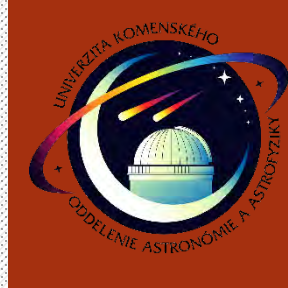
Photometry



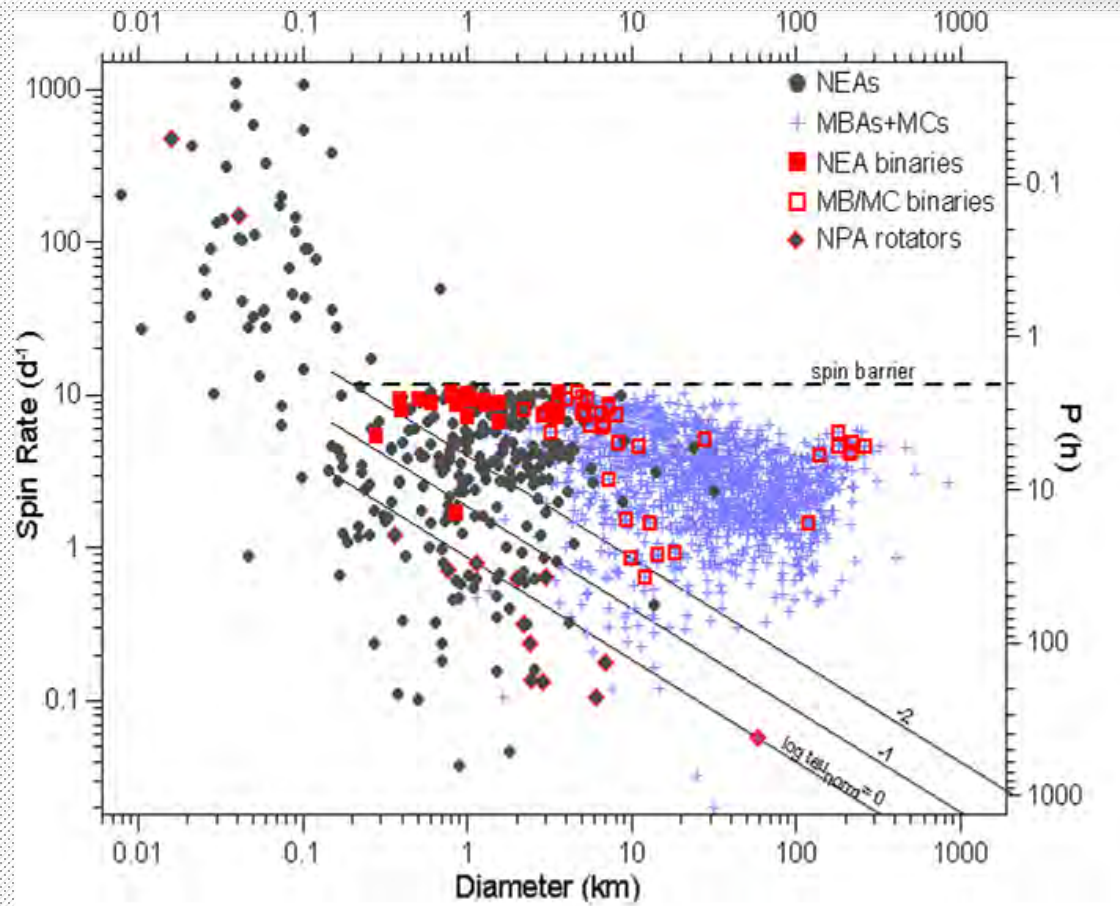
- 17. mag
- NEO, Mc, MBA
- Rotation period
- Precession



Photometry – binary asteroid



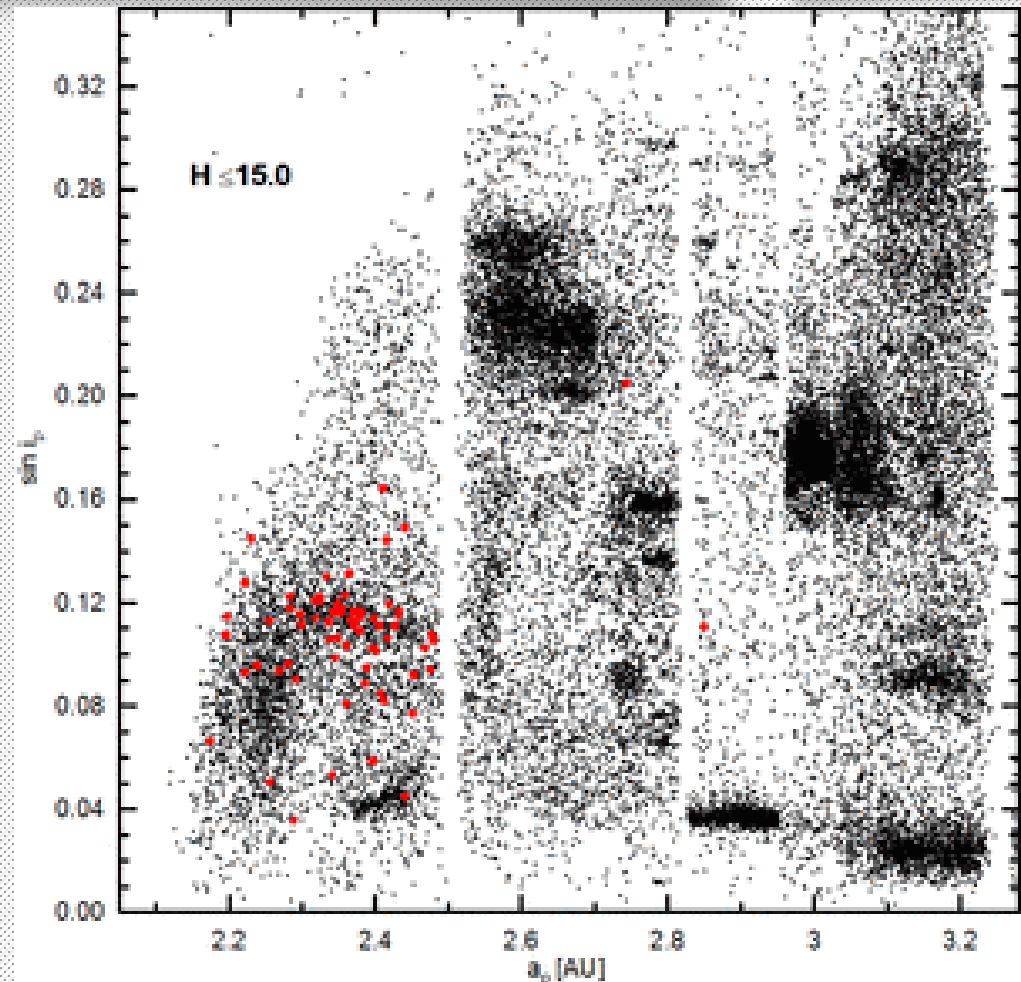
Rotational properties of asteroids spin barrier



Photometry – V-type asteroids



- asteroid Vesta
- Vesta family
- impact melting
- HED meteorites



Dynamical studies

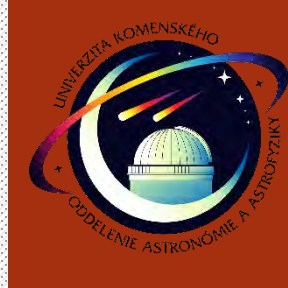
Bennu and Ryugu



Kováčová et al., Planet. Space Sci 185, 2020

Dynamical studies

Fast Lyapunov Indicator



8:3

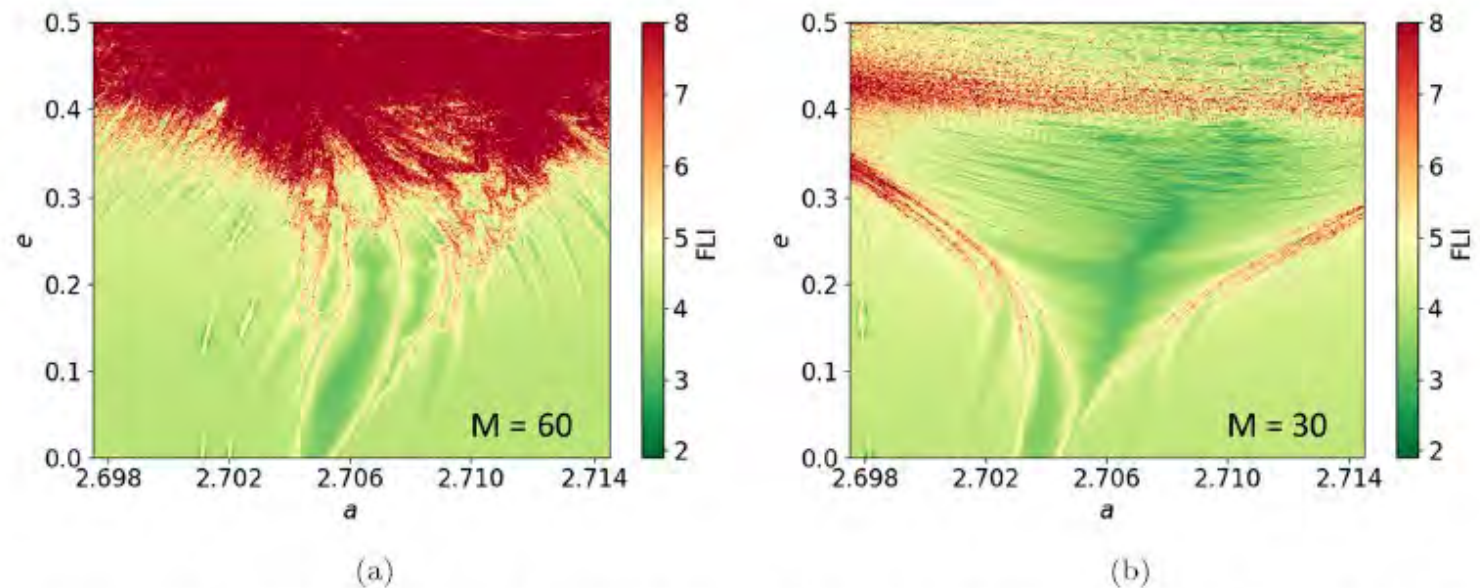
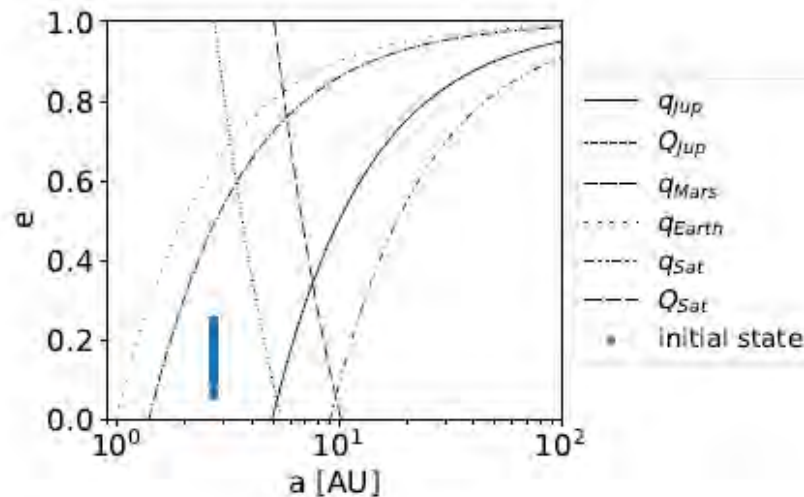
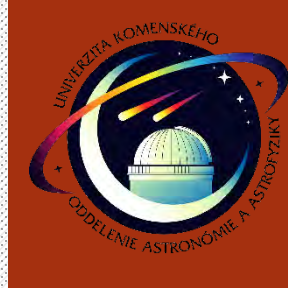
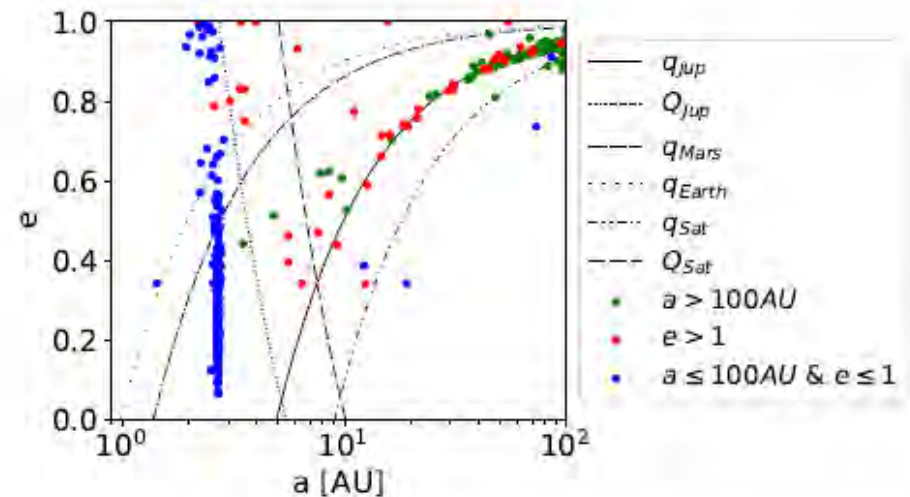


Figure 3. FLI maps of the 8:3 MMR with Jupiter in orbital plane of Ceres ($i = 10.594^\circ$, $\omega = 73.37^\circ$, $\Omega = 80.306^\circ$) for mean anomaly (a) $M = 60^\circ$ and (b) $M = 30^\circ$. Maps were computed for 5 kyr on a grid of 400×400 initial conditions (test particles) whereas semi-major axis $a \in \langle 2.6975, 2.7145 \rangle \text{ AU}$ and eccentricity $e \in \langle 0, 0.5 \rangle$. Colour scale represents FLI values determined as a maximum from values obtained for three different initial deviation vectors.

Dynamical studies



(a)



(b)

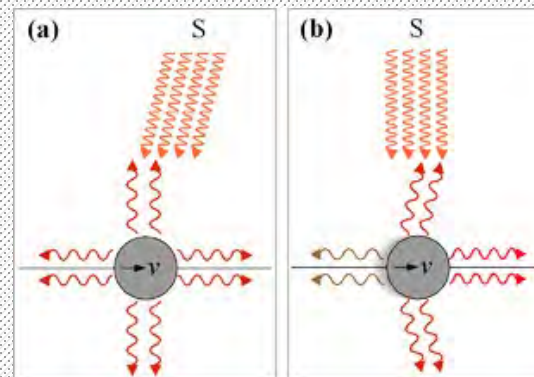
Figure 11. Initial (a) and end (b) state of simulation for unstable particles ($M = 60^\circ$). If particles reached the orbit with $a > 100 \text{ AU}$ (green dots) or hyperbolic orbit (red dots) on the graph there are their last known parameters displayed before they got ejected to these orbits. Black lines represent perihelion and aphelion distances of selected planets.

Non-gravitational effects



- Non-gravitational influence of the Sun:

- Solar radiation pressure
- Poynting-Robertson drag
- Solar wind



- Strongly influencing dynamical evolution of dust grains in the Solar System:

- L4 and L5 points under action of non-gravitational effects
- modelling of extrasolar dust discs

Galactic astrophysics

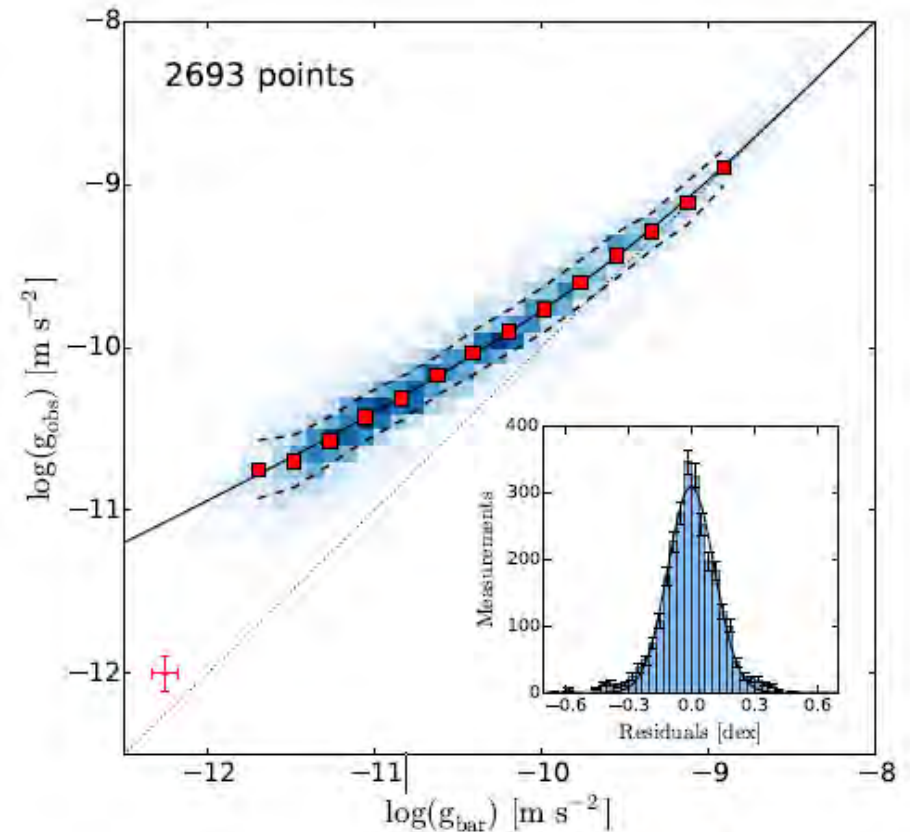


- kinematics and dynamics of the Galaxy:
 - GAIA – galactic flare
 - GAIA – galactic warp
 - simulation – comparison of MOND and NEWTON theories of gravity
- effects of galactic tides on the Solar System
- theory of gravity

Theory of gravity

$$g_{\text{obs}} = \mathcal{F}(g_{\text{bar}}) = \frac{g_{\text{bar}}}{1 - e^{-\sqrt{g_{\text{bar}}/g_{\dagger}}}}$$

- two body problem - not elliptical motion
- 10x reduced size of the Oort cloud





roman nagy

Galactic flare

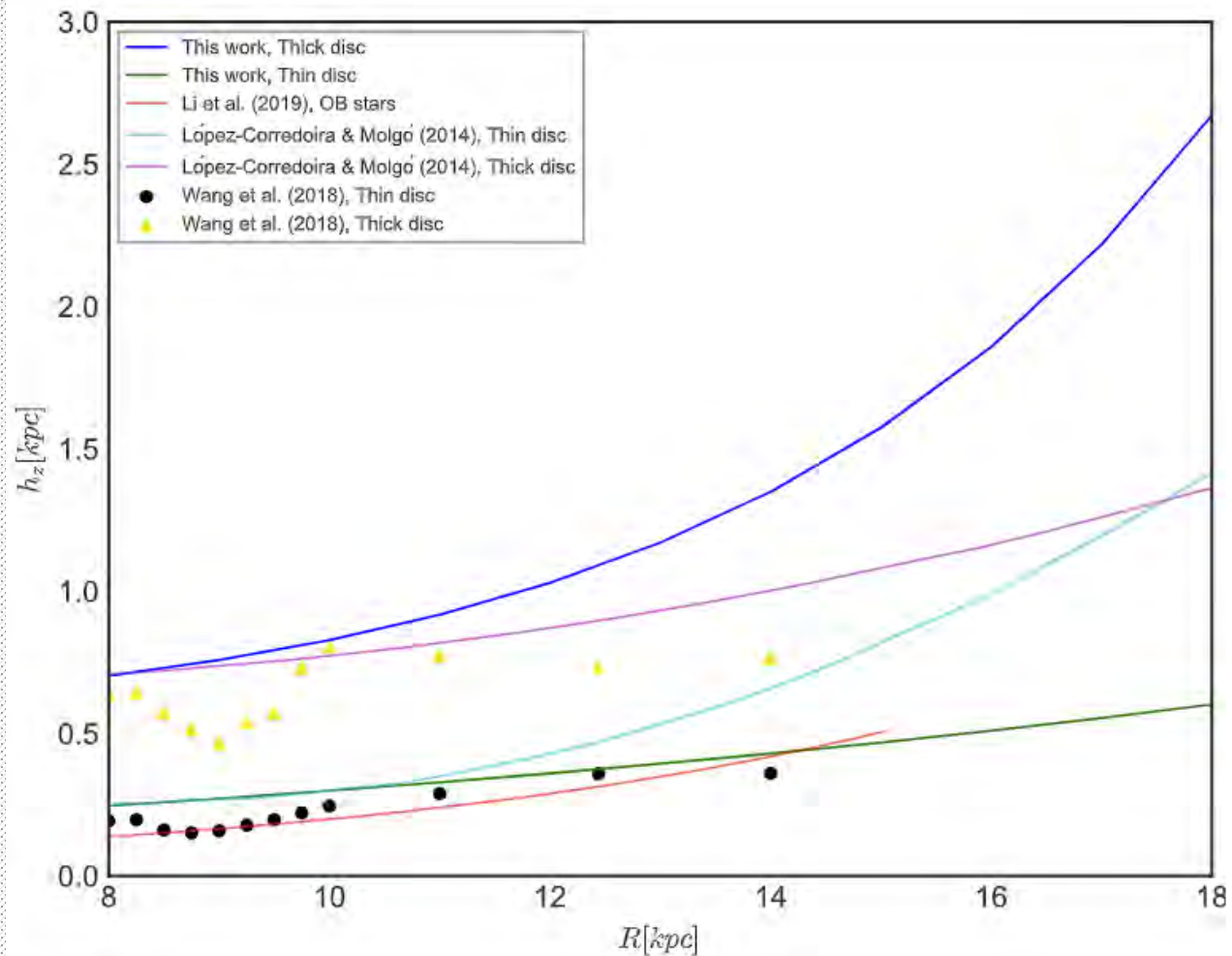


- Flare – thickening of the Galactic disc with Galactocentric distance R

$$\rho(R, z) = \rho_0(R) \times \exp(-R/h_r) \times \exp(-|z|/h_z) ; \quad h_z(R)$$

- Especially in remote regions of the Galaxy
- Observed in hydrogen gas as well as in stellar populations

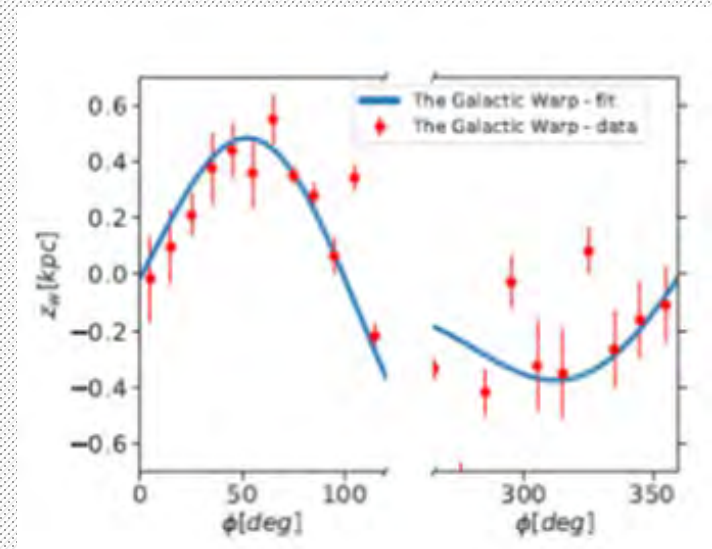
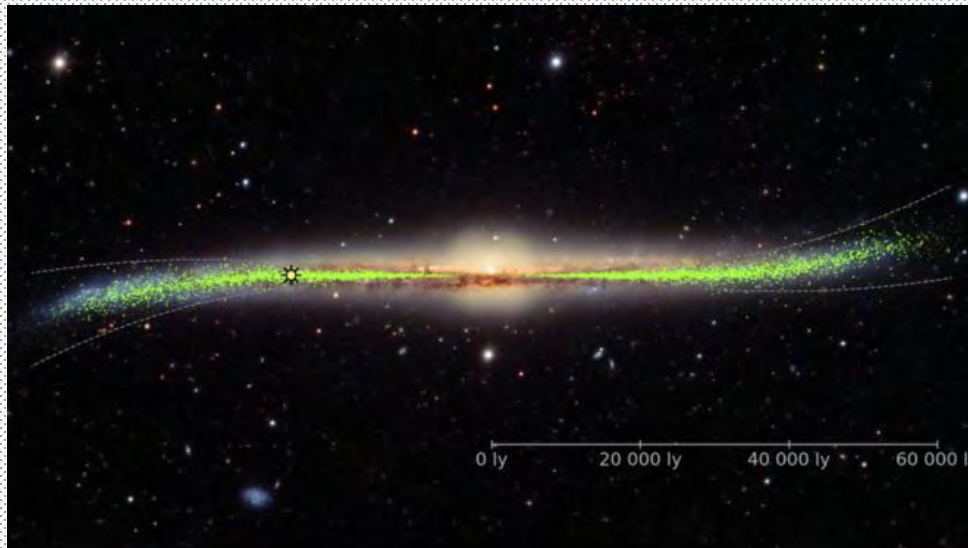
Galactic flare



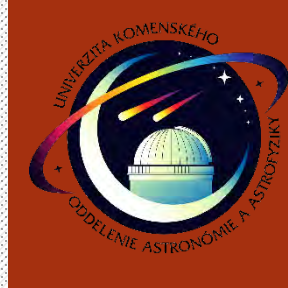
Galactic warp



- Warp – “S” shape warping of the Galactic disc



Space debris



- “All the man-made objects which are orbiting the Earth and have no further application.”

Satellite Astra



Upper stage Agena D



Al_2O_3



MLI



Probe Phobos Grunt



Upper stage Delta



Fragment SOCIT4 37



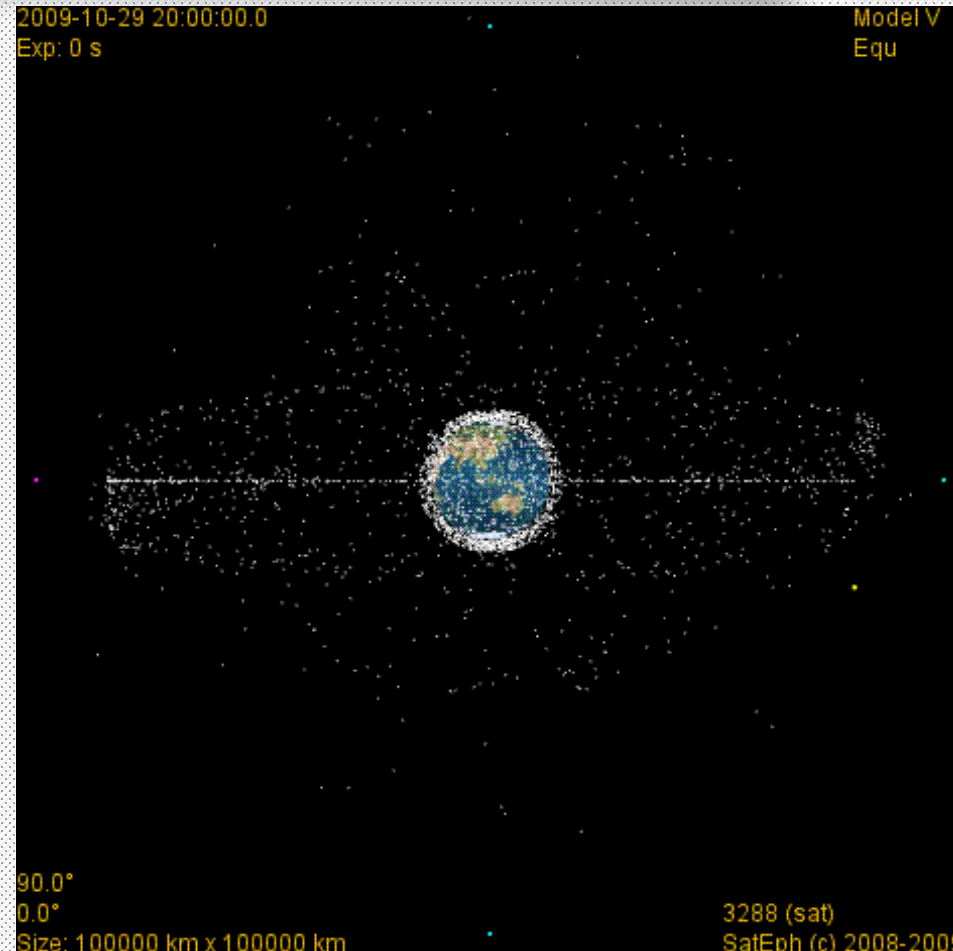
HST solar panel

70 cm telescope – space debris



2009-10-29 20:00:00.0
Exp: 0 s

Model V
Equ



FMPI AGO

70 cm Newton telescope



- Three major programs on AGO 70cm:
 - **Astrometry, surveys** - to maintain the objects catalogue, to discover new objects, cooperation with partners
 - **Photometry, light curves (LC)** - attitude state characterization, attitude motion determination, attitude evolution monitoring
 - **Photometry, colors and color indices (CI)** - surface properties, space-weathering aging monitoring
 - **Reflectance spectroscopy**

Optical observations, color photometry

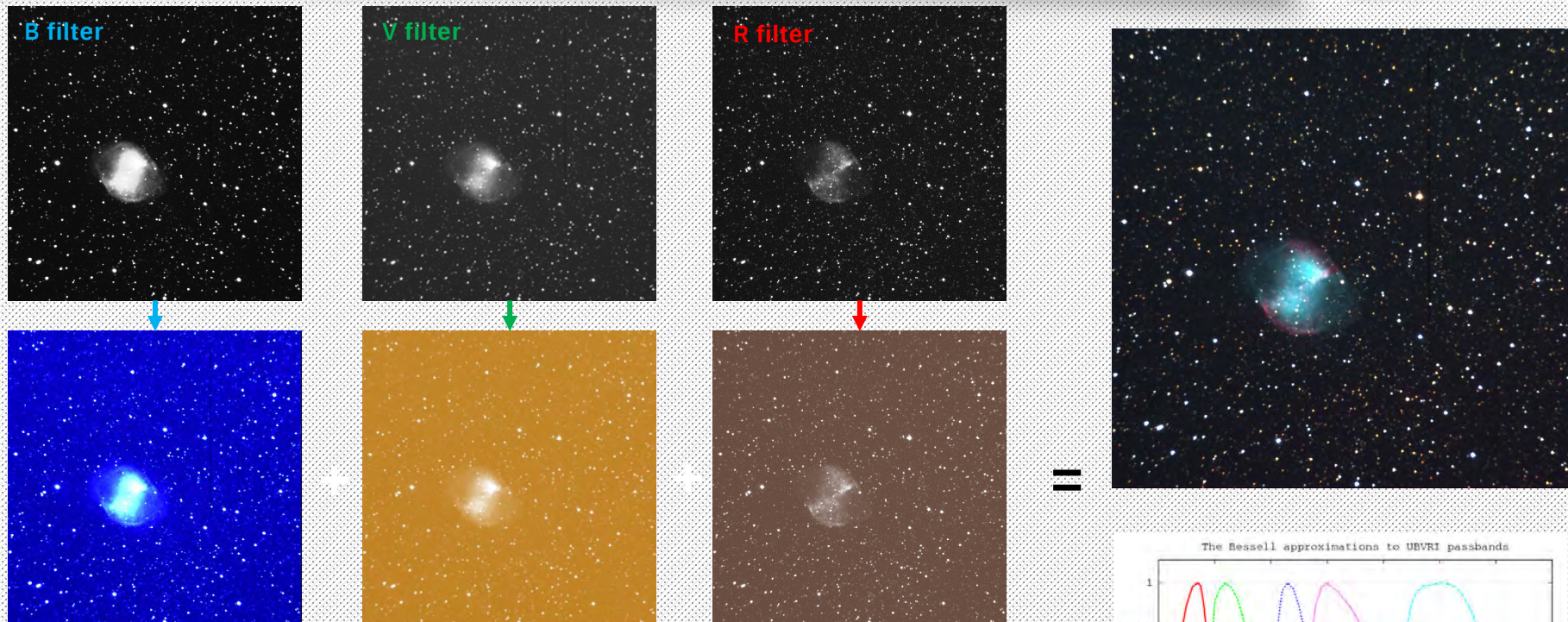
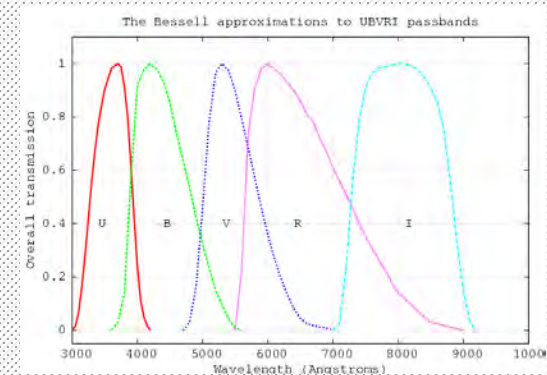
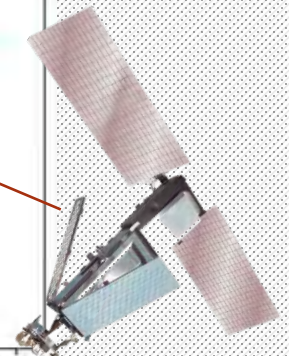
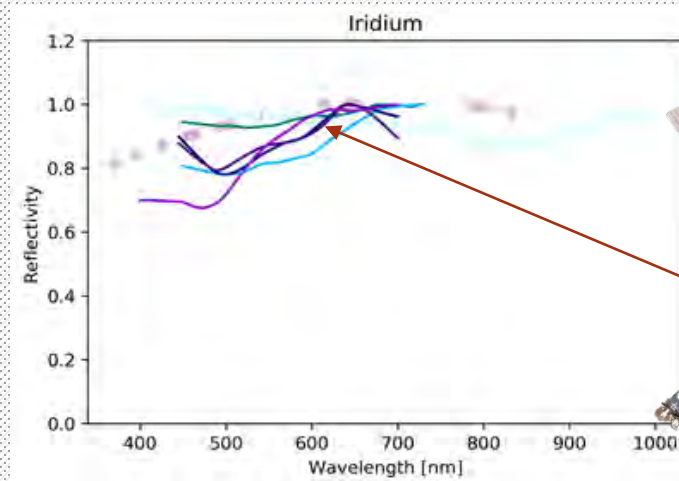
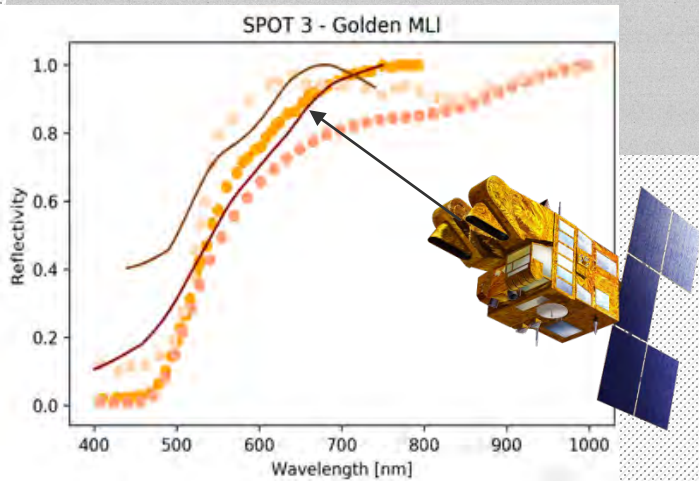
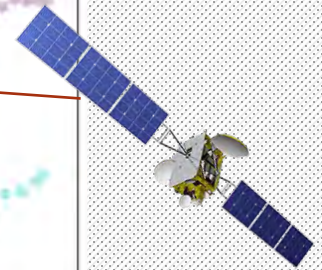
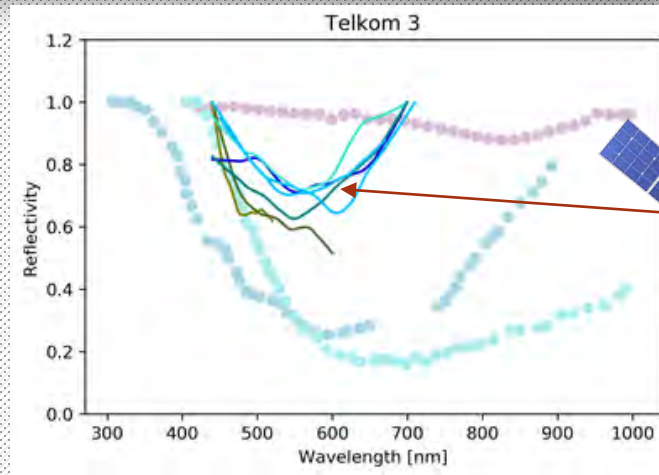
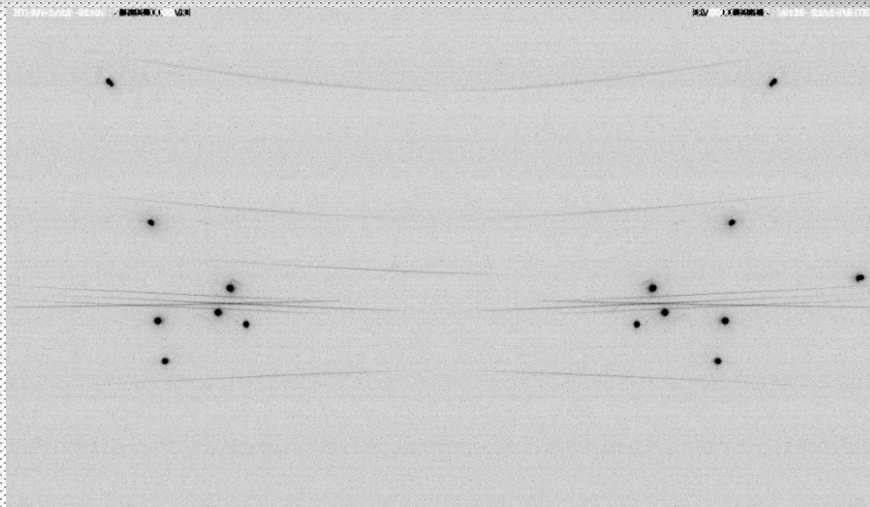
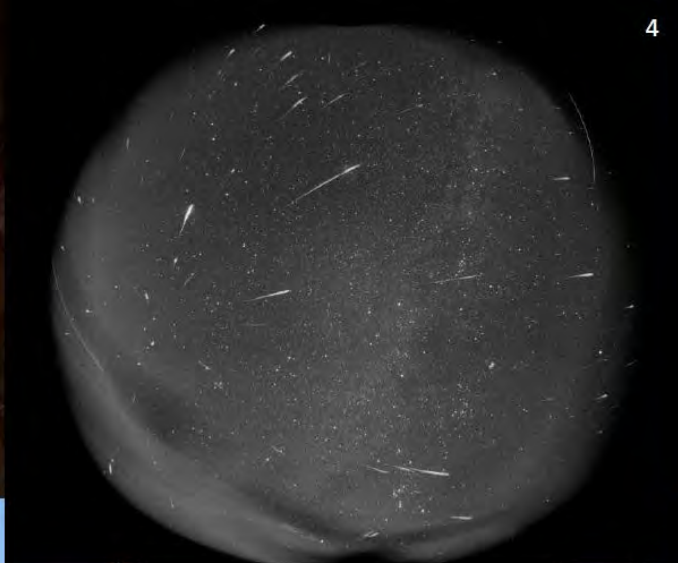


Figure – Dumbell nebula M27. Images acquired in B (left), V (middle) and R filters (right). Images acquired by AGO 70-cm telescope. Used exposure was 60s.



Space debris spectra by AMOS spectral cameras







Meteor astronomy



- observations:
 - video
 - radar
 - spectroscopic

Leonid meteor shower in 1998



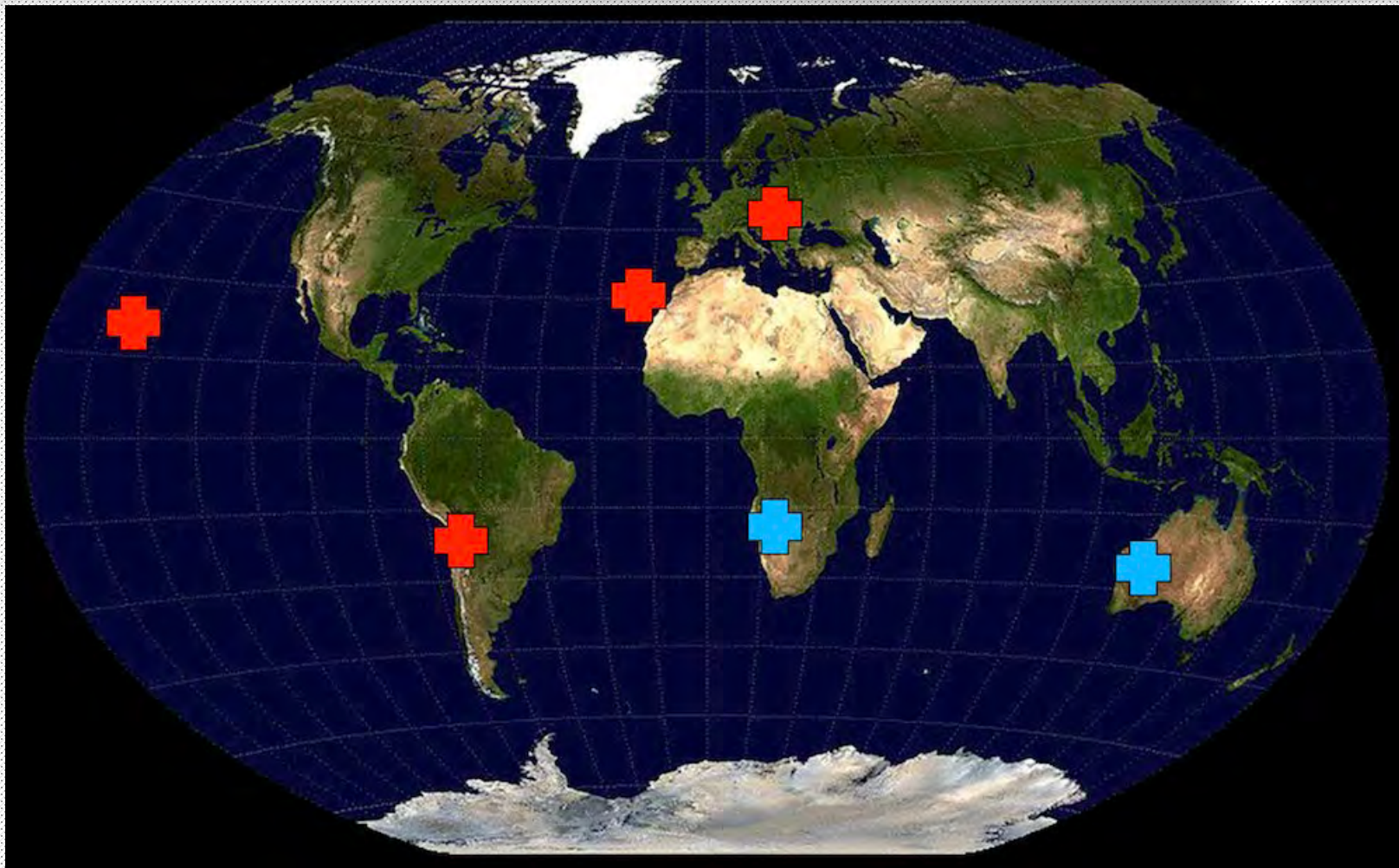
AGO Modra, FMFI UK

Video observations



- AMOS – All-sky Meteor Orbit System
- from 2009
- 10 stations (Slovakia, Canary Islands, Chile, Hawaii)

AMOS



AMOS – Teide Observatory



AMOS – Hawaii

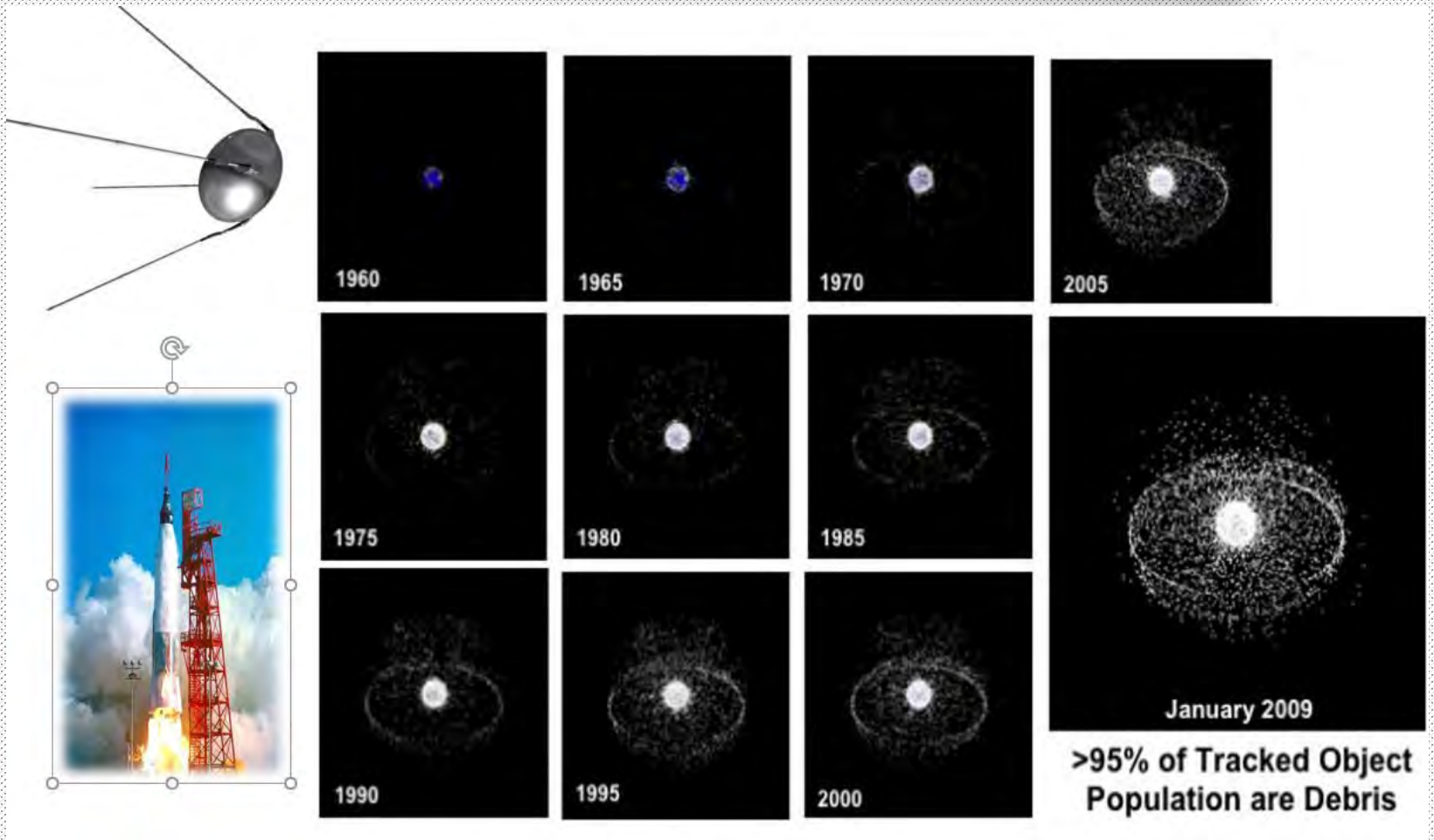


One thousand Geminids above Tenerife Dec.13/14, 2017





Space debris, history



FMPI AGO

70 cm Newton telescope

